Introduction Background

Space-based remote sensing of solar radiation scattered by hydrogen (H) atoms is often used to infer the density of Earth's uppermost atmosphere based on an assumed functional form for the global density distribution. Here, we present a alternative approach to conventional parametric estimation that is based instead on tomographic inversion of satellite observations of scattered ultraviolet emission at 121.6 nm.

Forward Model

Atmostpheric H emission intensity y, as measured from planetocentric position r along a line-of-sight (LOS) in the direction $\hat{\boldsymbol{n}}$, is related to the unknown H density ρ in terms of distance $l \equiv |\mathbf{r}' - \mathbf{r}|$ (measured in R_E , radius of earth) along LOS as follows:

$$y(\mathbf{r}, \widehat{\mathbf{n}}) \propto \int_{0}^{l_{\infty}(\widehat{\mathbf{n}})} \rho(\mathbf{r}') dl$$
 (1)

Utilizing standard basis functions in polar coordinates, the 2-D plane is discretized into polar rectangles, and assuming the density distribution is constant within each polar rectangle, the following can be derived:

$$= \mathbf{L}\mathbf{x} + \mathbf{w} \qquad (2$$

- $\mathbf{y}, \mathbf{w} \in \mathbb{R}^{I}$, I observed emission intensity and Poisson distributed additive noise
- $\mathbf{x} \in \mathbb{R}^{J}$, H density in J polar rectangle
- $\mathbf{L} \in \mathbb{R}^{I \times J}$, observation matrix constructed from discretizing (1). L is a function of \hat{n} , r



2D Demonstration of variables defined above in a discretized planetecentric coordinate. Satellite has position vector *r* with look angle \hat{n} and LOS (yellow)

Aim

the nature of limited data and errors in Due to measurement, the inverse problem is ill-conditioned. Previous techniques include:

- Tikhonov: $J(x, s) = ||y Lx||^2 + |\alpha_1^2||Dx||^2$ Data Fidelity Regularizer Linear optimization problem
- Over smooth solution, edges not preserved
- Total Variation: $J(\mathbf{x}, \mathbf{s}) = ||\mathbf{y} \mathbf{L}\mathbf{x}||^2 + |\alpha_1^2|\mathbf{D}\mathbf{x}|^2$

Data Fidelity Regularizer

Well known for edge preservation Edges are not limited, can appear randomly Cost function is non-differentiable

TOMOGRAPHIC RECONSTRUCTION OF ATMOSPHERIC DENSITY WITH MUMFORD-SHAH FUNCTIONALS David Ren, Lara Waldrop, Farzad Kamalabadi

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Method Mumford-Shah Functionals

$$J(\boldsymbol{u}) = \int_{\Omega} (\boldsymbol{u} - \boldsymbol{f})^2 d\boldsymbol{A} + \alpha_1^2 \int_{\Omega \setminus \Gamma} |\nabla \boldsymbol{u}|^2 d\boldsymbol{A} + \alpha_2^2 |\Gamma|^2$$

- Ω, Γ, u, f : Image domain, boundary, and fields
- Originally used for Image Segmentation
- Produces **piece-wise smooth** images
- The functional is non-differentiable

Ambrosio–Tortorelli Approximation

$$J(\mathbf{x}, \mathbf{s}) = \underbrace{\left| |\mathbf{y} - L\mathbf{x}| \right|^{2}}_{Data \ Fidelity} + \underbrace{\alpha_{1}^{2} ||\mathbf{D}\mathbf{x}||_{W_{s}}^{2} + \alpha_{2}^{2} ||\mathbf{D}\mathbf{s}||_{2}^{2} + \alpha_{3}^{2} ||\mathbf{s}||_{2}^{2}}_{Regularizer}$$

- **D**: Discrete differencing matrix
- *s*: Edge field corresponding to each "pixel"
- W_s: Weighting matrix according to edge field s
- Allowed for Maximum-*a-posteriori* estimation interpretation
- Non-linear optimization problem due to W_s

Optimization Steps

Step 1: Initialization

- Initial guess x_0, s_0
- Compute $W_s = diag(1 s^2)$



Step 3: Convergent Solution Smooth edge preservation

Model independent structure of H density

Observations along LOS that transit very near the earth (blue) or through its shadow (grey) are neglected due to complicated scattering physics occurring there that invalidates the forward model

Variatio

Mumford-Shah

Results **Synthetic Model**



Viewing Geometries



Illustration of Line-of-sight coverage of the 2-D field (green) from the measurement positions along the satellite trajectory (red)

Reconstruction Results



Spherical Harmonics (2D)

Realistic polar 10% - 20%

10

0

 $\rho\left(r_k, \theta = \frac{\pi}{2}, \phi\right)$

 $+ B_{lm}(r_k) \sin(m\phi) Y_{lm}(\theta)$

depletion region

• Radial Res.: $1.5 R_E$

• Angular Res.: $\frac{\pi}{15}$





Summary & Future Work

1. We proposed the technique of tomographic **inversion** to make sense of the H density distribution, and the result is model-independent. 2. We proposed using Mumford-Shah functionals as the regularizer, which allows "smart" edge preservation while maintaining **smoothness**. It also produces superior results comparing to Total Variation method. 3. We will aim at developing optimal regularization. parameter selection methods and investigating the fundamental limits of reconstruction resolution.



In addition, the reconstruction quality of a single "pixel" depends on:

error in the reconstructions using MS functionals for the linear CASE A viewing geometry as a function of sampling frequency

Relative reconstructions using MS functionals for elliptical CASE B viewing geometry as a of sampling function frequency

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