

## Introduction

In this paper, a reweighted sparse low-rank nonnegative tensor factorization (RSLRNTF) method is proposed to restore an HSI. It takes an HSI as a third-order tensor and factorizes it into the combination of a few component tensors where each one is the outer product of a low-rank matrix (coding matrix) and a vector (atom). A reweighted  $L_1$  norm is added to coding matrices to enforce their sparsity. The low-rankness in both spatial and spectral domain is in line with the spatial and spectral correlation in an HSI. Furthermore, we add nonnegativity constraint to both coding coefficients matrices and dictionary to learn parts-based representation of HSI, which facilitates preserving local structure information.

## RSLRNTF

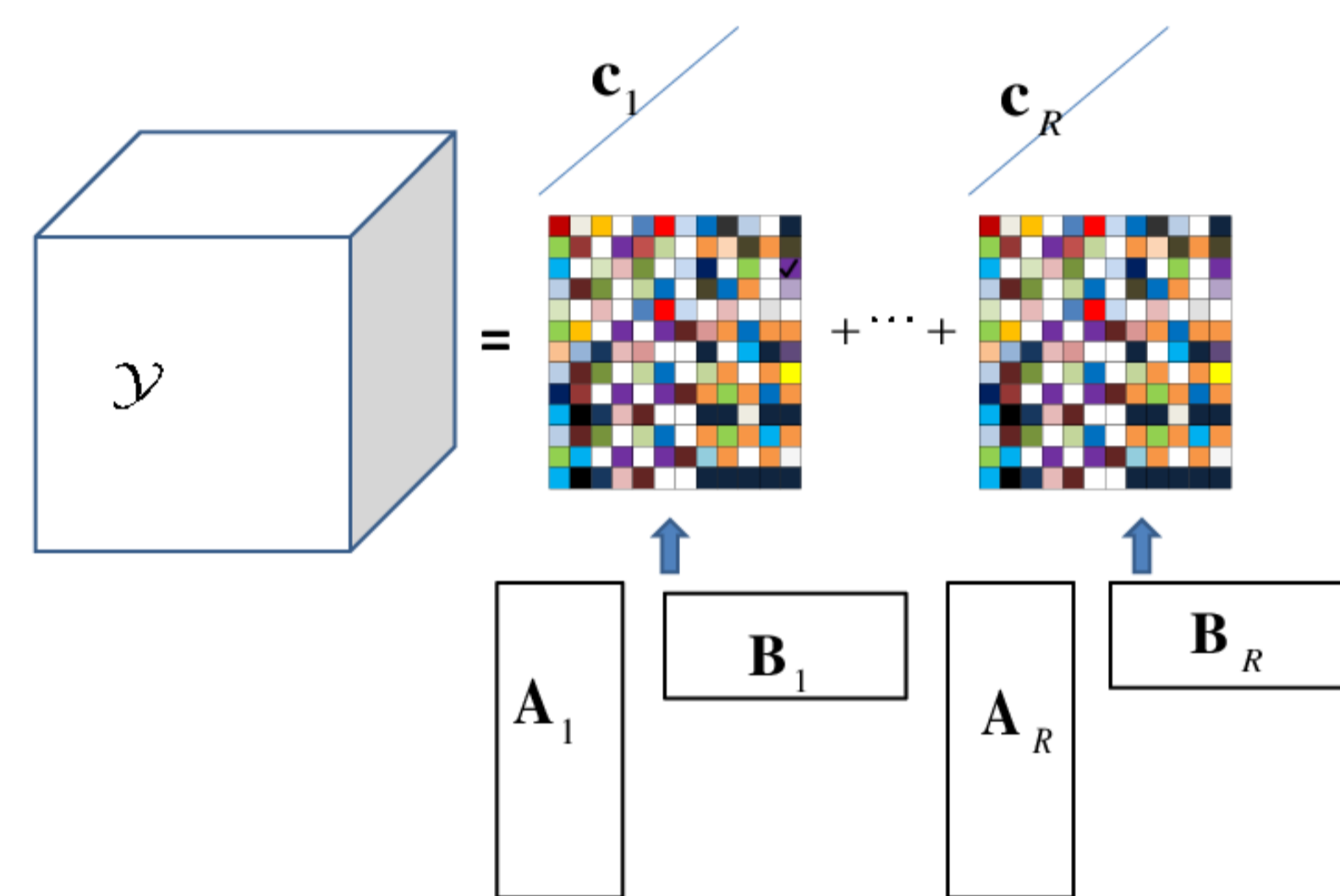


Figure 1: Framework of proposed method.

$$\min f(\mathbf{A}, \mathbf{B}, \mathbf{C}) = \|\mathcal{Y} - \sum_{r=1}^R \mathbf{A}_r \mathbf{B}_r^T \circ \mathbf{c}_r\|_F^2 + \frac{\delta}{2} \|\mathbf{A} \mathbf{B}^T - \mathbf{1}_{\mathbf{I} \times \mathbf{J}}\|_F^2 + \lambda \sum_{r=1}^R \|\mathbf{W}_r \odot (\mathbf{A}_r \mathbf{B}_r^T)\|_1 \quad s.t. \quad \mathbf{A}, \mathbf{B}, \mathbf{C} \geq 0$$

where,  $\mathbf{A} = [\mathbf{A}_1 \cdots \mathbf{A}_R]$ ,  $\mathbf{B} = [\mathbf{B}_1 \cdots \mathbf{B}_R]$ ,  $\mathbf{C} = [\mathbf{c}_1 \cdots \mathbf{c}_R]$  and  $\odot$  is element-wise product.

## Highlights

- ▶  $\mathbf{A}_r \mathbf{B}_r^T$ : low-rank coding matrix derives **spatial low-rankness**.
- ▶  $\mathbf{C}$ : low-rank dictionary represents **spectral low-rankness**.
- ▶  $\|\mathbf{W}_r \odot (\mathbf{A}_r \mathbf{B}_r^T)\|_1$ : reweighted  $L_1$  norm makes coding matrix **sparser**.
- ▶  $\frac{\delta}{2} \|\mathbf{A} \mathbf{B}^T - \mathbf{1}_{\mathbf{I} \times \mathbf{J}}\|_F^2$ : avoid trivial solutions.
- ▶ **Nonnegativity** constraints promote local details preserving.
- ▶ **Tensor** model preserves all the information in an HSI.

## Update rules

$$\begin{aligned} \mathbf{A} &\leftarrow \mathbf{A} * (\mathbf{Y}_{(1)}^T \mathbf{M} + \delta \mathbf{1}_{\mathbf{I} \times \mathbf{J}} \mathbf{B} + \mu \mathbf{U}) ./ (\mathbf{A} \mathbf{M}^T \mathbf{M} + \delta \mathbf{A} \mathbf{B}^T \mathbf{B} + \mu \mathbf{A}) \\ \mathbf{B} &\leftarrow \mathbf{B} * (\mathbf{Y}_{(2)}^T \mathbf{M} + \delta \mathbf{1}_{\mathbf{I} \times \mathbf{J}}^T \mathbf{A} + \mu \mathbf{V}) ./ (\mathbf{B} \mathbf{M}^T \mathbf{M} + \delta \mathbf{B} \mathbf{A}^T \mathbf{A} + \mu \mathbf{B}) \\ \mathbf{C} &\leftarrow \mathbf{C} * (\mathbf{Y}_{(3)}^T \mathbf{M}) ./ (\mathbf{C} \mathbf{M}^T \mathbf{M}) \\ \mathbf{U}_r &\leftarrow \mathbf{U}_r * (\mu \mathbf{A}_r) ./ (\mu \mathbf{U}_r + \lambda \mathbf{W}_r \mathbf{V}_r) \\ \mathbf{V}_r &\leftarrow \mathbf{V}_r * (\mu \mathbf{B}_r) ./ (\lambda \mathbf{W}_r^T \mathbf{U}_r + \mu \mathbf{V}_r) \\ \mathbf{W}_r &\leftarrow 1 ./ (\mathbf{U}_r \mathbf{V}_r^T + \epsilon) \end{aligned}$$

## Experimental results on simulated data

Table 1: MPSNR of Different Methods on Simulated Data.

Noise variance $\sigma \in [0, 0.05]$	$\sigma \in [0, 0.1]$	$\sigma \in [0, 0.15]$	
Original	34.49	26.89	24.84
LRMR [1]	43.12	39.12	36.63
LRTV [2]	43.04	37.63	36.26
PARAFAC [3]	38.58	37.51	35.99
TDL [4]	38.99	36.61	31.75
LRTDTV [5]	42.63	39.14	36.94
<b>RSLRNTF</b>	<b>43.87</b>	<b>40.01</b>	<b>37.21</b>

## Visual comparison on simulated data

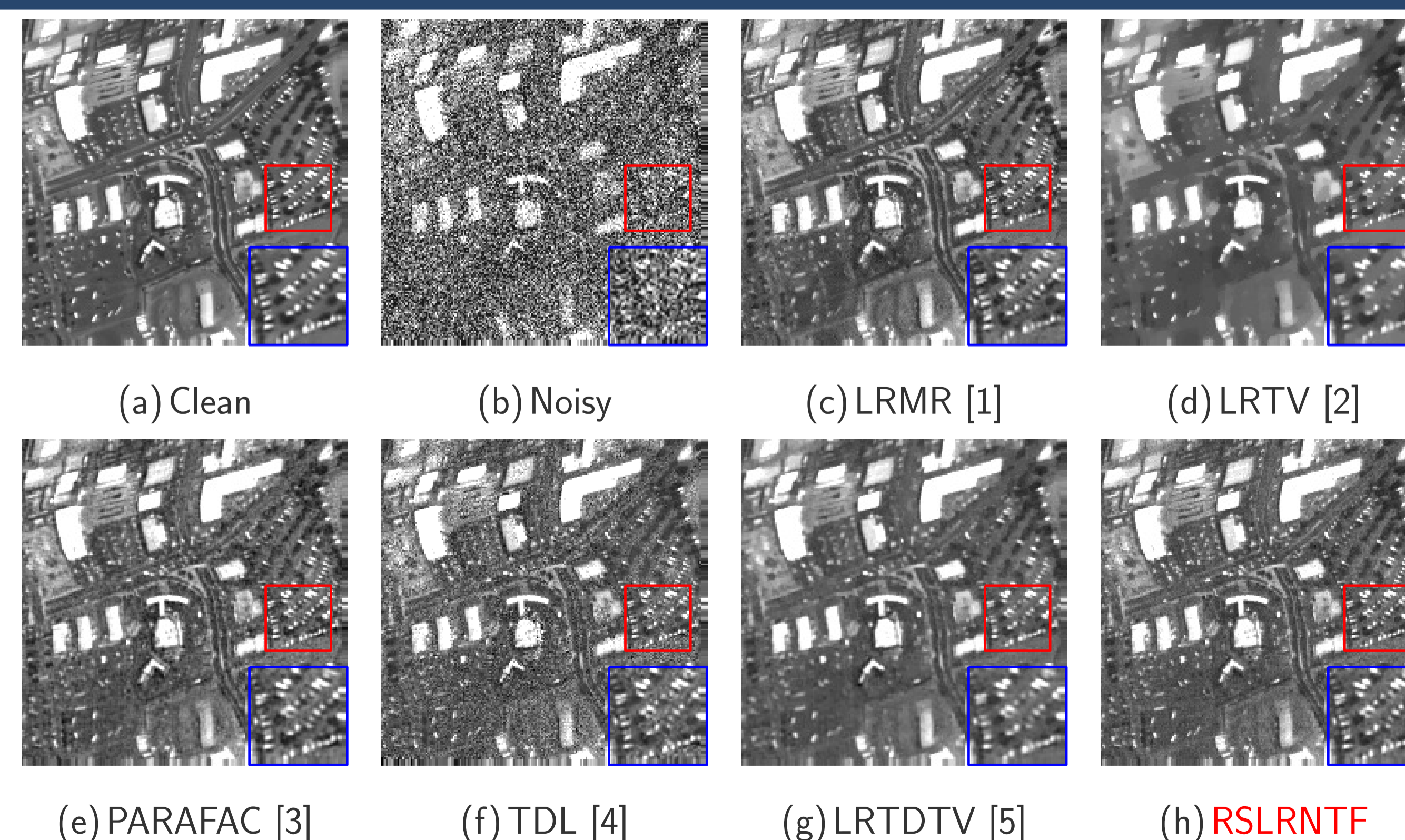


Figure 2: Denoising results of band 25 in simulated data when  $\sigma \in [0 - 0.1]$ .

## Experimental results on real-world data

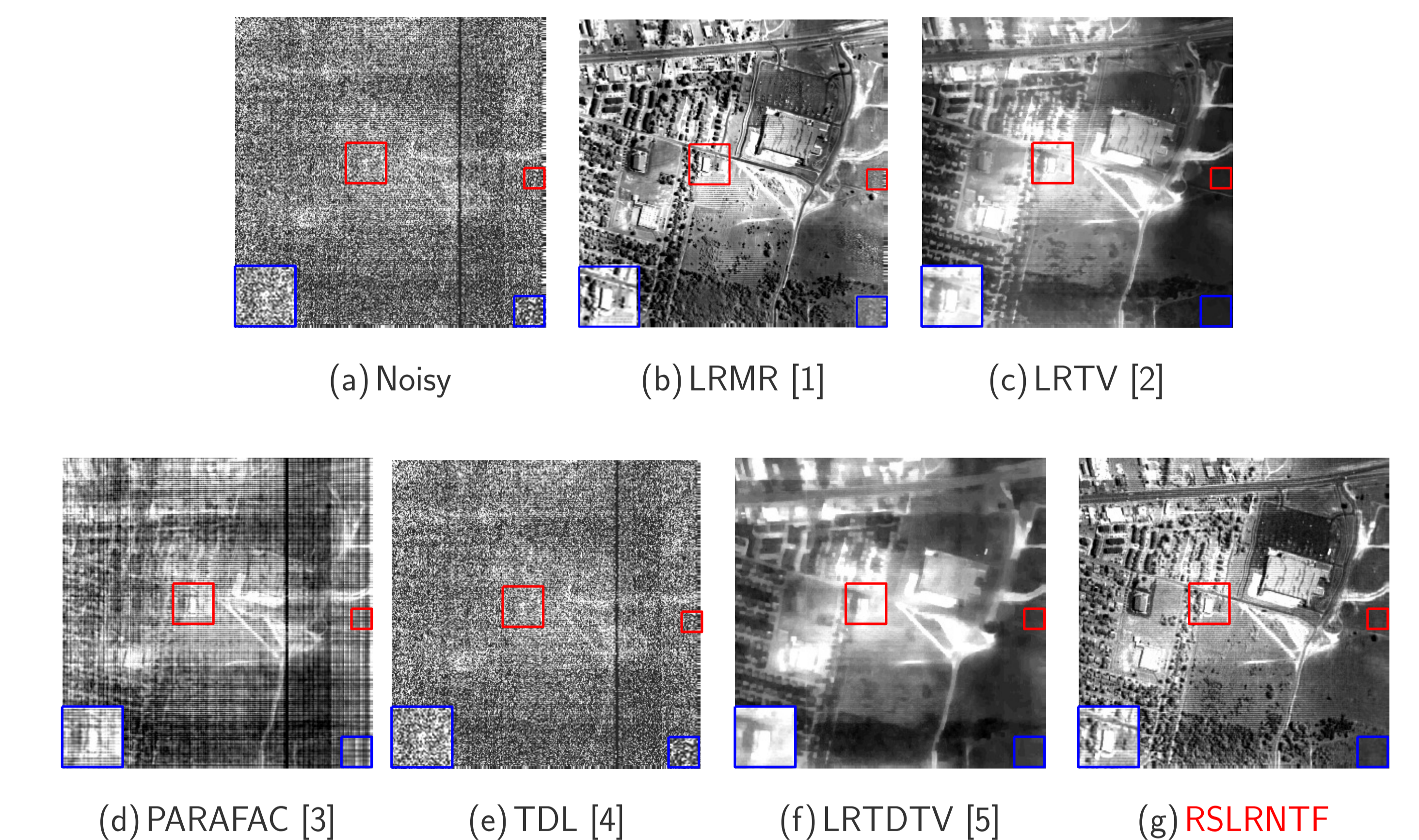


Figure 3: Denoising results of band 208 in the Urban data set.

## References

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- [3] X. Liu, S. Bourennane, and C. Fossati, "Denoising of hyperspectral images using the parafac model and statistical performance analysis," *IEEE Trans. Geosci. Remote Sens.*, vol. 50, no. 10, pp. 3717–3724, Oct. 2012.
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- [5] Y. Wang, J. Peng, Q. Zhao, Y. Leung, X. Zhao, and D. Meng, "Hyperspectral image restoration via total variation regularized low-rank tensor decomposition," *IEEE J. Sel. Topics Appl. Earth Observ. Remote Sens.*, vol. 11, no. 4, pp. 1227–1243, 2018.