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Persistent Hyperspectral Observations of the Urban Lightscape

J. Baur, G. Dobler, F. Bianco, S. Koonin, M. Sharma, A. Karpf Urban Observatory

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Center for Urban Science + Progress

CUSP Urban Observatory

Better cities through imaging

Main Focus

Energy

electrical energy consumption patterns

• stresses on the power grid

Environment

- measuring the impact of cities on air quality
- detecting plumes

Public Health & Policy

light pollution circadian rhythm https://cuspuo.org
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Andreas Karpf

Project Manager

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Our Team







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Project Manager

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IDEAS grant AR0000886-F0217:

Grid Dynamics and Energy Consumption Patterns through remote observations of city lights

Imaging the NYC Lightscape

Remote, Persistent, Synoptic





Instrumentation



Middleton Spectral Vision

vertical slit (30μ m) aperture spectrograph

Imperx B1621 camera (mono) KAI-02050 CCD image sensor

> 5.5 μm 8.98 mm x 6.78 mm 1600 x 1200 pixels 42 fps



848 spectral channels

deployment setup



Instrumentation



Middleton Spectral Vision

vertical slit (30μ m) aperture spectrograph

Imperx B1621 camera (mono) KAI-02050 CCD image sensor

5.5 μm 8.98 mm x 6.78 mm sensor format: 1600 x 1200 pixels 42 fps







Instrumentation





1600 vertical pixels

/pan direction



Instrumentation







1600 vertical pixels

pan direction

Instrumentation



Middleton Spectral Vision

Specim V10E vertical slit (30µm) aperture spectrograph

> 0.4 - 1.1 μm 0.72 x 10⁻³μm (fwhm)

> > Imperx B1621 camera (mono) KAI-02050 CCD image sensor

pixel size:5.5 μmsensor format:8.98 mm x 6.78 mmresolution:1600 x 1200 pixelsframe rate:42 fps







Cleaned Data



Hyperspectral Survey

Manhattan, 2013

"A Hyperspectral Survey of New York City Lighting Technology" Dobler, Ghandehari, Koonin & Sharma Sensors, vol.16 no12, 2016 DOI:<u>10.3390/s16122047</u>

Manhattan view (north facing)



New York City Lighting Technologies





wavelength [range: 0.4-1.0 microns]



Hyperspectral Survey Brooklyn, 2017

Brooklyn view (south facing)



Lamp Spectral Power D is tribution Database

courtesy of Johanne Roby (CEGEP Sherbrook, QC) www.lspdd.com

- lab-measured spectral distributions of 254 light types
- BLACK-Comet spectrometer (StellarNet)



0.4 0.5 0.6 0.7 0.8 0.9 wavelength: λ / μm



Identifying Sources



integrated intensity [log scaled]

Source Detection

Selection

training sample: 500 source, 500 noise





1.0

1.0

Sample Split



333 testing set

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Supervised Classifier Training



Classification



Classification



Performance



333 testing set

1	Test on sa) tests, shuffle	
		848 features (all λ)	2 features (VIS - NIR)
	accuracy score	97.4 % (± 0.7)	96.2 % (± 0.9)
	false ⊕	0.5 % (± 0.4)	0.7 % (± 0.4)
	false ⊖	2.1 % (± 0.7)	3.1 % (± 0.9)



Performance



Test on samples100 tests, shuffle			
		848 features (all λ)	2 features (VIS - NIR)
	accuracy score	97.4 % (± 0.7)	96.2 % (± 0.9)
	false ⊕	0.5 % (± 0.4)	0.7 % (± 0.4)
	false ⊖	2.1 % (± 0.7)	3.1 % (± 0.9)

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2 Labelling full cube 5.1 x10⁶ pixels cpu time (1 core) 0 960 sec 0 60 sec

Transferability



1	Test on sa	mples 100) tests, shuffle
		0.40 fa atuma a	0 fa atuma a

	848 features (all λ)	2 features (VIS - NIR)
accuracy score	97.4 % (± 0.7) 96.2 % (± 0.4)	96.2 % (± 0.9) 93.8 % (± 0.9)
false ⊕	0.5 % (± 0.4) 0.4 % (± 0.2)	0.7 % (± 0.4) 0.2 % (± 0.1)
false ⊖	2.1 % (± 0.7) 3.3 % (± 0.5)	3.1 % (± 0.9) 6.0 % (± 0.6)

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2	Labelling full cube		2.5 x1	0 ⁶ pixels
		_		

cpu time	∿ 960 sec	∿ 60 sec
(1 core)	∿ 460 sec	∿ 29 sec

Transferability



1	Test on sa	mples
		848 fe

100 tests,	shuffle
------------	---------

	848 features (all λ)	2 features (VIS - NIR)
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cpu time	∿ 960 sec	∿ 60 sec
(1 core)	∿ 460 sec	∿ 29 sec

Next Step Technology Penetration & Time Series









Thank You

for your attention





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