

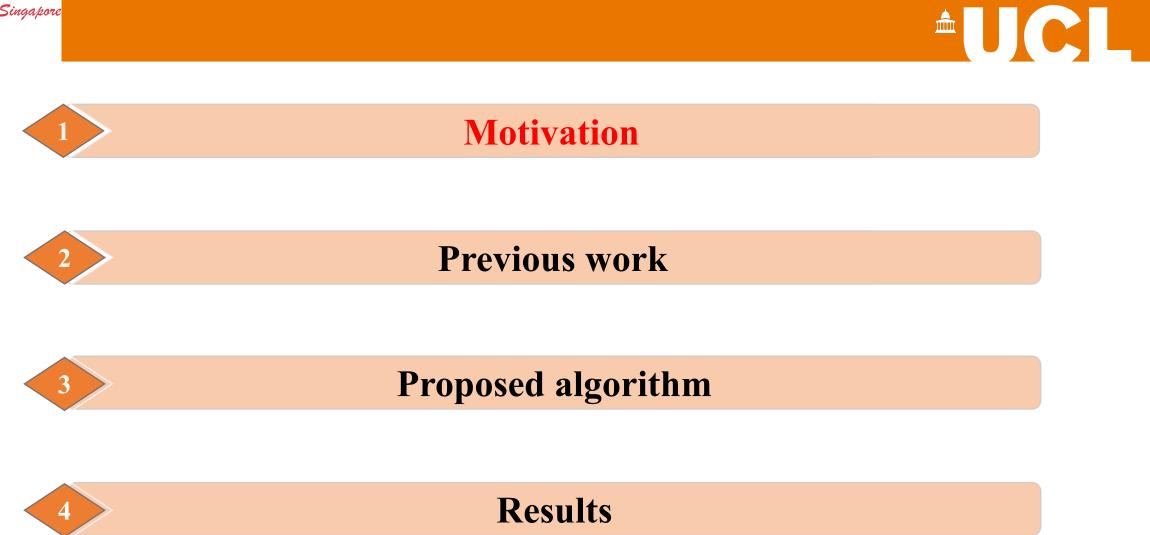


# BLIND UNMIXING USING A DOUBLE DEEP IMAGE PRIOR

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#### ✤ HSI linear unmixing model

Y = EA + N

#### where,

(Y, E, A) is (HSI reflectance spectrum, material reflectance /endmember signature matrix, abundance matrix);

N is generally additive Gaussian noise;

### ✤ Goal

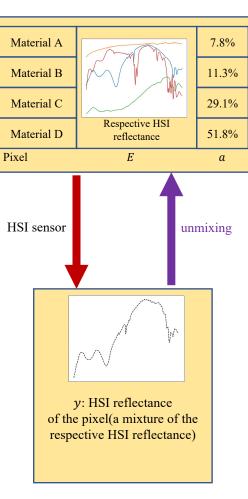
Given *Y*, estimate *A* and *E*.

$$\hat{E}, \hat{A} = \arg\min_{E,A} \frac{1}{2} \|Y - EA\|_{F}^{2} + R(A)$$
  
s.t.,  $E \ge 0, A \ge 0, A^{T} \mathbf{1}_{r} = \mathbf{1}_{n}$ 

### Previous methods

- Most network based blind unmixing methods cannot guarantee to generate physically meaningful unmixing results due to the lack of effective guidance<sup>[1]</sup>.
- > The performance of most unmixing networks with training guidance is limited by the quality of the guidance.
- ✤ Motivation
- Can we propose a new network that can surpass the performance limitation imposed by training guidance?

[1] D. Hong et al., "Endmember-Guided Unmixing Network (EGU-Net): A General Deep Learning Framework for Self-Supervised Hyperspectral Unmixing," in IEEE Transactions on Neural Networks and Learning Systems, doi: 10.1109/TNNLS.2021.3082289.



### Fig.1 HSI unmixing Problem

(1)









# Unmixing using Deep Image Prior (UnDIP) **\*** DIP<sup>[2]</sup>

#### Problem setting

Consider the image inverse problem, such as denoising, given by:

$$x^* = \underset{x}{\operatorname{argmin}} \|x - x_0\|_2^2 + R(x) \quad (1)$$

where,

 $x_0$  is the noisy image;

R is a regularizer on x.

> <u>DIP technique</u>

DIP propose to solve it by:

$$\theta^* = \underset{x}{\operatorname{argmin}} \|f_{\theta}(z) - x_0\|_2^2 \quad (2)$$

where,

 $f_{\theta}(z)$  is a neural network parameterized by  $\theta$  with a random input z.

≻ <u>Note</u>

♣After learning, the network  $f_{\theta}$  can replace the regularizer R.
♣The restored image is given by  $x^* = f_{\theta^*}(z)$ .

[2] Dmitry Ulyanov, Andrea Vedaldi, and Victor Lempitsky, "Deep image prior," International Journal of Computer Vision, vol. 128, no. 7, pp. 1867–1888, Mar 2020.



✤ UnDIP<sup>[3]</sup>

# Unmixing using Deep Image Prior (UnDIP)

Suppose some existing methods give  $\vec{E}$ ,

- Cons:
- No E itself;
- Limited by  $\widehat{E}$ .

given  $\tilde{E}$  $\widehat{E}$ ,  $\widehat{A}$  $\hat{A} = \operatorname{argmin}_{A} \frac{1}{2} \| Y - \tilde{E}A \|_{F}^{2} + R(A)$ (4)  $= \overline{\operatorname{argmin}_{E,A} \frac{1}{2} \|Y - EA\|_F^2 + R(A)}$ (3)  $s.t.A \ge 0, A^T 1_r = 1_n$  $s.t., E \ge 0, A \ge 0, A^T \mathbf{1}_r = \mathbf{1}_n$ DIP  $\theta^* = \operatorname{argmin}_{\theta} \frac{1}{2} \left\| Y - \tilde{E} f_{\theta}(z) \right\|_{F}^{2}$ (5) $\hat{A} = f_{A^*}(z)$ Fig. 2. Proposed convolutional network architecture with one skip connection. This network is used as  $f_{\theta}$  for UnDIP in the experiments. Different layers in the network are shown with specific colors.

[3] Rasti, B., Koirala, B., Scheunders, P., & Ghamisi, P. (2021). UnDIP: Hyperspectral Unmixing Using Deep Image Prior. *IEEE Transactions on Geoscience and Remote Sensing*, 1–15. https://doi.org/10.1109/TGRS.2021.3067802









# Endmember estimation via EDIP

Like UnDIP, suppose an estimation of abundance  $\tilde{A}$  is given by existing methods, then (1) would reduce to endmember estimation problem:

$$\hat{E} = \arg\min_{E} \frac{1}{2} \left\| Y - E\tilde{A} \right\|_{F}^{2}$$
  
s.t.,  $E \ge 0$ 

According to DIP technique, we propose to solve it via

$$\hat{\theta}_E = \min_{\theta_E} \frac{1}{2} \left\| Y - f_{\theta_E}(z_E) \tilde{A} \right\|_F^2$$

Where,  $f_{\theta_E}(z_E)$  is the network with learnable parameters  $\theta_E$  and random input  $z_E$   $\circ$ 

- ▶ Last layer of  $f_{\theta_E}$  is sigmoid to impose  $E \ge 0$ .
- > Endmember given by  $\hat{E} = f_{\hat{\theta}_E}(z_E)$ .





## Abundance estimation via EDIP

Similarly, suppose an estimation of endmember  $\tilde{E}$  is given by existing methods, then (1) would reduce to abundance estimation problem:

$$\hat{A} = \arg\min_{E} \frac{1}{2} \left\| Y - \tilde{E}A \right\|_{F}^{2}$$
  
s.t.,  $A \ge 0, A^{T} \mathbf{1}_{F} = \mathbf{1}_{n}$ 

According to DIP technique, we propose to solve it via

$$\hat{\theta}_A = \min_{\theta_A} \frac{1}{2} \left\| Y - \tilde{E} f_{\theta_A}(z_A) \right\|_F^2$$

Where,  $f_{\theta_A}(z_A)$  is the network with learnable parameters  $\theta_A$  and random input  $z_{A\circ}$ 

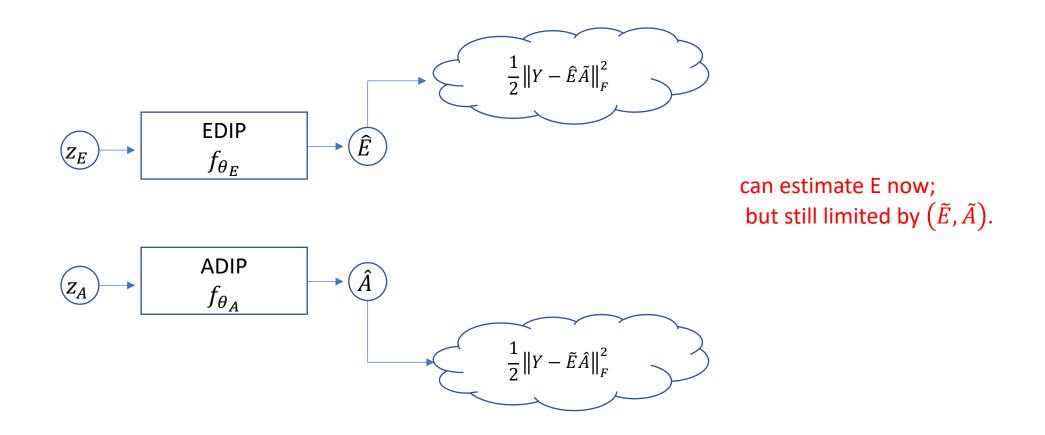
- ► Last layer of  $f_{\theta_A}$  is softmax to impose  $A \ge 0$ ,  $A^T \mathbf{1}_r = \mathbf{1}_n$ .
- → Abundance given by  $\hat{A} = f_{\hat{\theta}_A}(z_A)$ .





# Unmixing using Double DIP (BUDDIP)

EDIP + ADIP

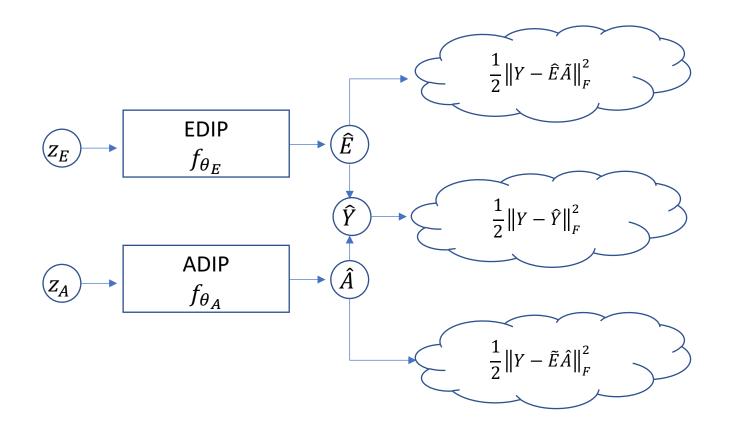






# Unmixing using Double DIP (BUDDIP)

EDIP + ADIP



✓ We also propose the third loss: Blind unmixing loss:  $L_{BU} = \frac{1}{2} \|Y - \hat{Y}\|_{F}^{2}$ 

✓ Final loss  $L = \alpha_1 L_{EDIP} + \alpha_2 L_{ADIP} + \alpha_3 L_{BU}$ 

✓ New interpretation Based on the guidance  $(\tilde{E}, \tilde{A})$ , find a better solution.









### Experiment on Synthetic Data

data	generate according to [4].		
competitors	UnDIP <sup>[3]</sup> ; SiVM <sup>[5]</sup> +FCLS <sup>[6]</sup> .		
metrics for endmember	average Spectral Angle Distance (aSAD) $aSAD(E, \hat{E}) = \frac{1}{r} \sum_{i=1}^{r} \frac{180}{\pi} \cos^{-1}(e_i, \hat{e}_i)$		
metrics for abundance	average Root mean square error (aRMSE) $aRMSE(A, \hat{A}) = \frac{1}{n} \sum_{i=1}^{n} \sqrt{\frac{1}{r}   a_i - \hat{a}_i  _2^2}$		
optimizer	ADAM		
Learning rate	1e-4		
epochs	4500		
hyper-parameters	$\alpha_1 = 0.1, \alpha_2 = 0.01, \alpha_3 = 1.0$		

[4] S. Jia and Y. Qian, "Constrained nonnegative matrix factorization for hyperspectral unmixing," IEEE Trans. Geosci. Remote Sens., vol. 47, no. 1, pp. 161–173, Jan. 2009.

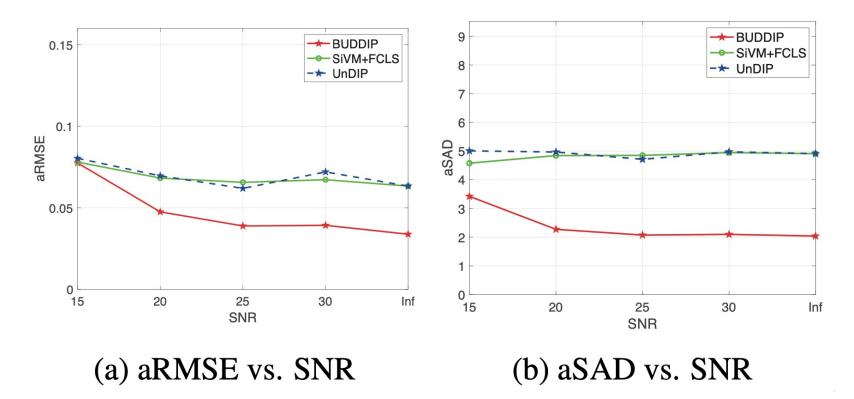
[5] Rob Heylen, Dz<sup>\*</sup>evdet Burazerovic, and Paul Scheun- ders, "Fully constrained least squares spectral unmix- ing by simplex projection," *IEEE Transactions on Geo- science and Remote Sensing*, vol. 49, no. 11, pp. 4112–4122, 2011.

[6] D. C. Heinz and Chein-I-Chang, "Fully constrained least squares linear spectral mixture analysis method for material quantification in hyperspectral imagery," IEEE Transactions on Geoscience and Remote Sensing, vol. 39, no. 3, pp. 529–545, 2001.



## Performance vs. SNR

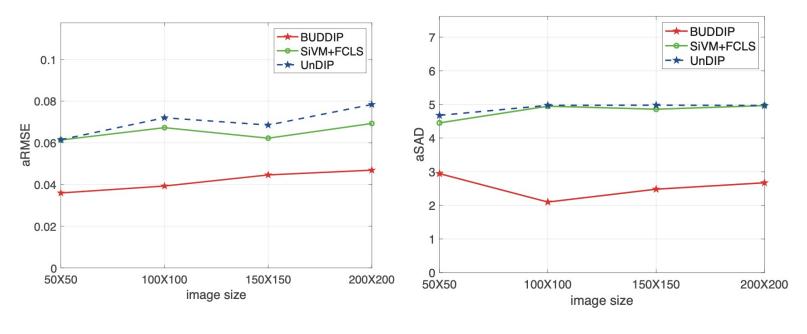
### We use default setting except SNR vary in [15,20,25,30, *inf*] dB.





### Performance vs. HSI image size

We use default setting except training size vary in  $\{50 \times 50, 100 \times 100, 150 \times 150, 200 \times 200\}$  pixels.



(a) aRMSE vs. image size

(b) aSAD vs. image size





### **Experiment Real Data**

#### Jasper Ridge<sup>[1]</sup>

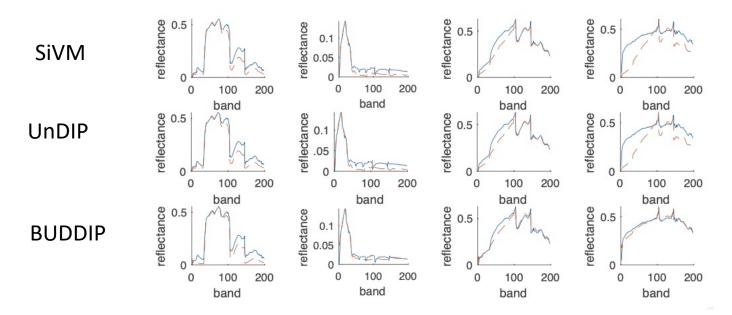


term	value
pixels	100 x 100
channels	198
endmembers	Road, Soil, Water, Tree



### Performance: Endmember estimation

We use default setting except  $\alpha_1 = \alpha_2 = \alpha_3 = 1.0$  and epoch=24000.



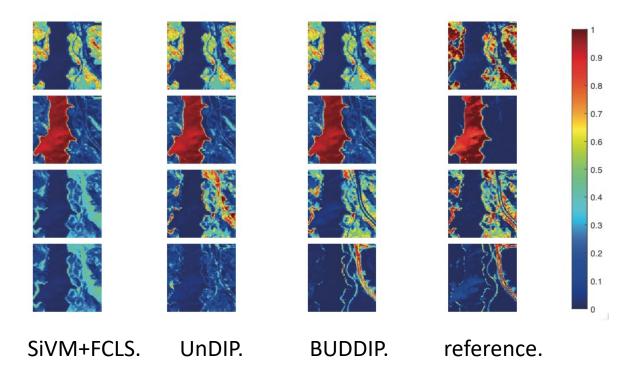
	SiVM +FCLS	UnDIP	BUDDIP
aSAD	11.349	11.349	6.8489

Solid blue line is true value, while dot line is estimated value.



## Performance: Abundance estimation

We use default setting except  $\alpha_1 = \alpha_2 = \alpha_3 = 1.0$  and epoch=24000.



	SiVM +FCLS	UnDIP	BUDDIP
aRMSE	0.1480	0.1748	0.1023



